

# Seeking the (very) Elusive

## Optimising your equipment, eyes and techniques

Can a new (personal) regime  
significantly improve your deep sky  
observing?

Keith Venables FRAS  
IoA Cambridge  
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# The basics for seeing the 'elusive'

## 'Find'

Locate (with certainty) the correct star field at a suitable magnification

“with the necessary magnification, locate the object position”

## 'See'

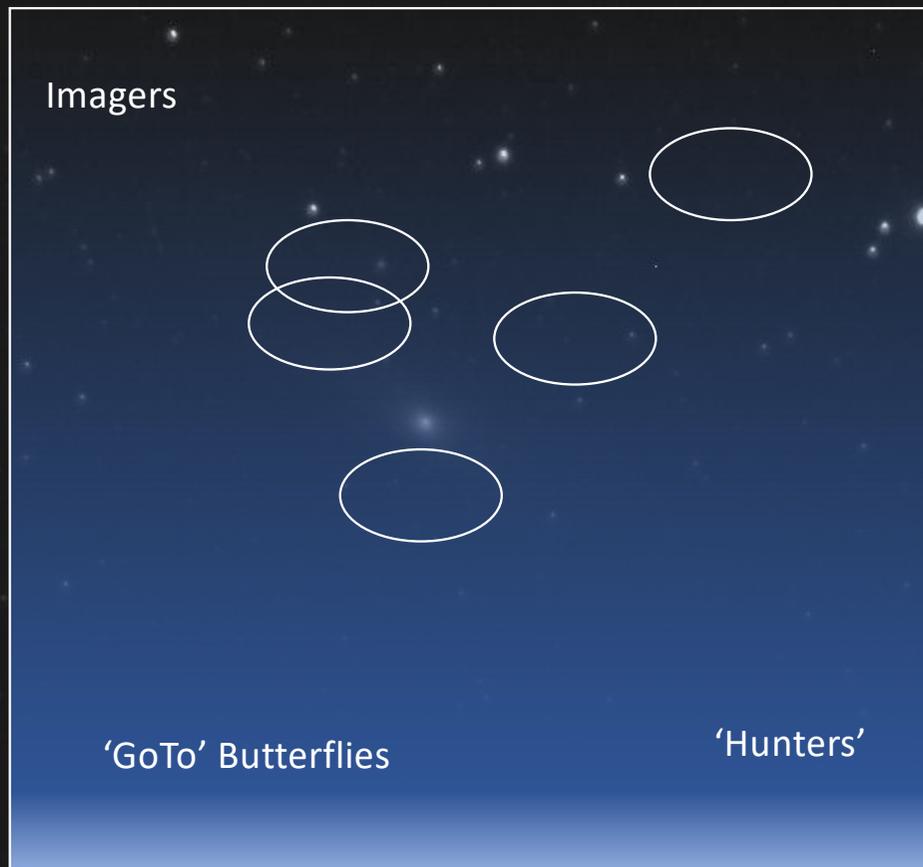
See the target, usually starting with a 'glimpse'. Be able to re-look and acquire almost immediately.

## 'Look'

Get comfortable and look for detail. Compare with reference images. Try different eyepieces and filters.....

# A spectrum of observer types

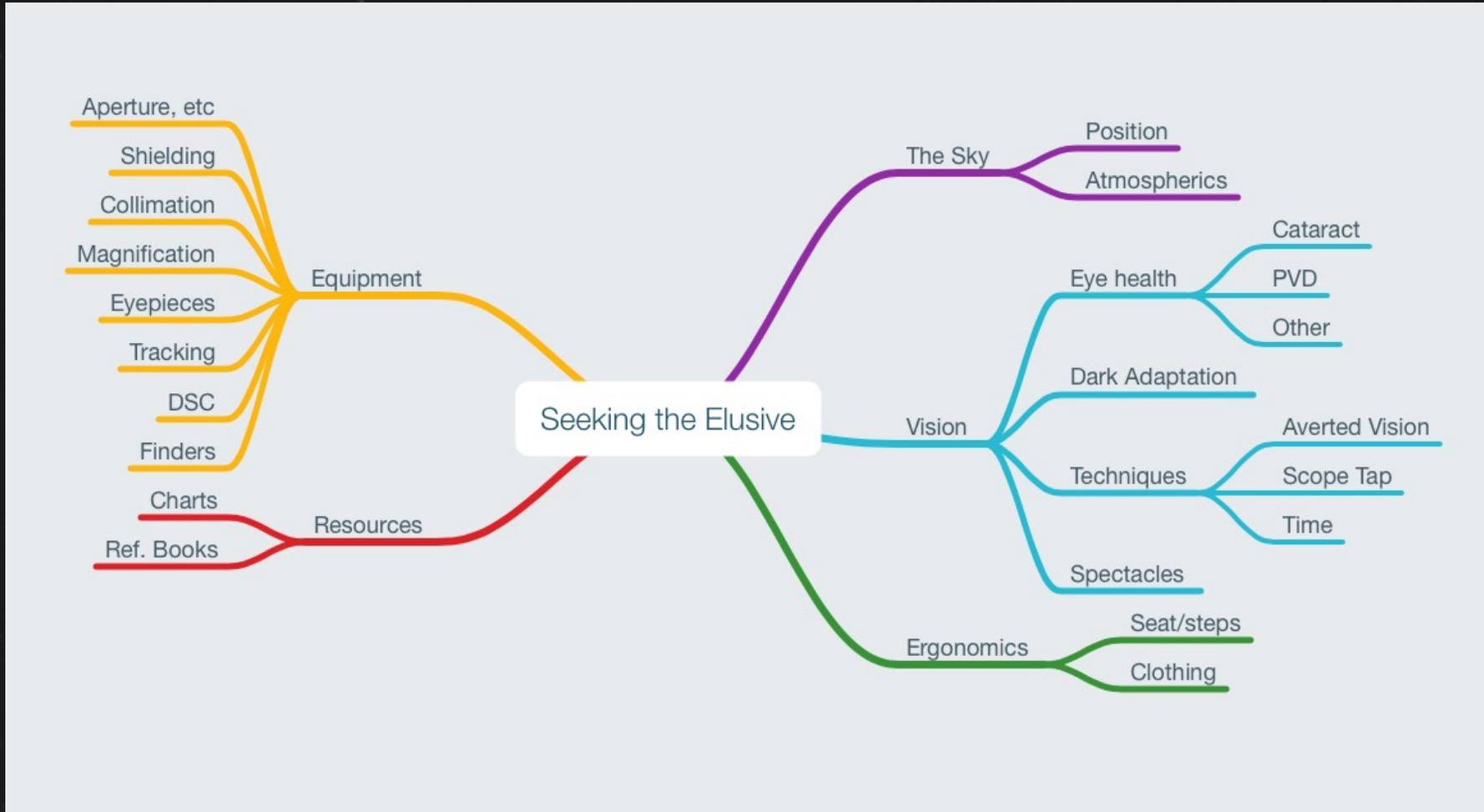
Time on Target



Time acquiring



# Exploring the factors



# Survey results considered

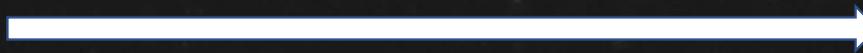
High reward



Low reward



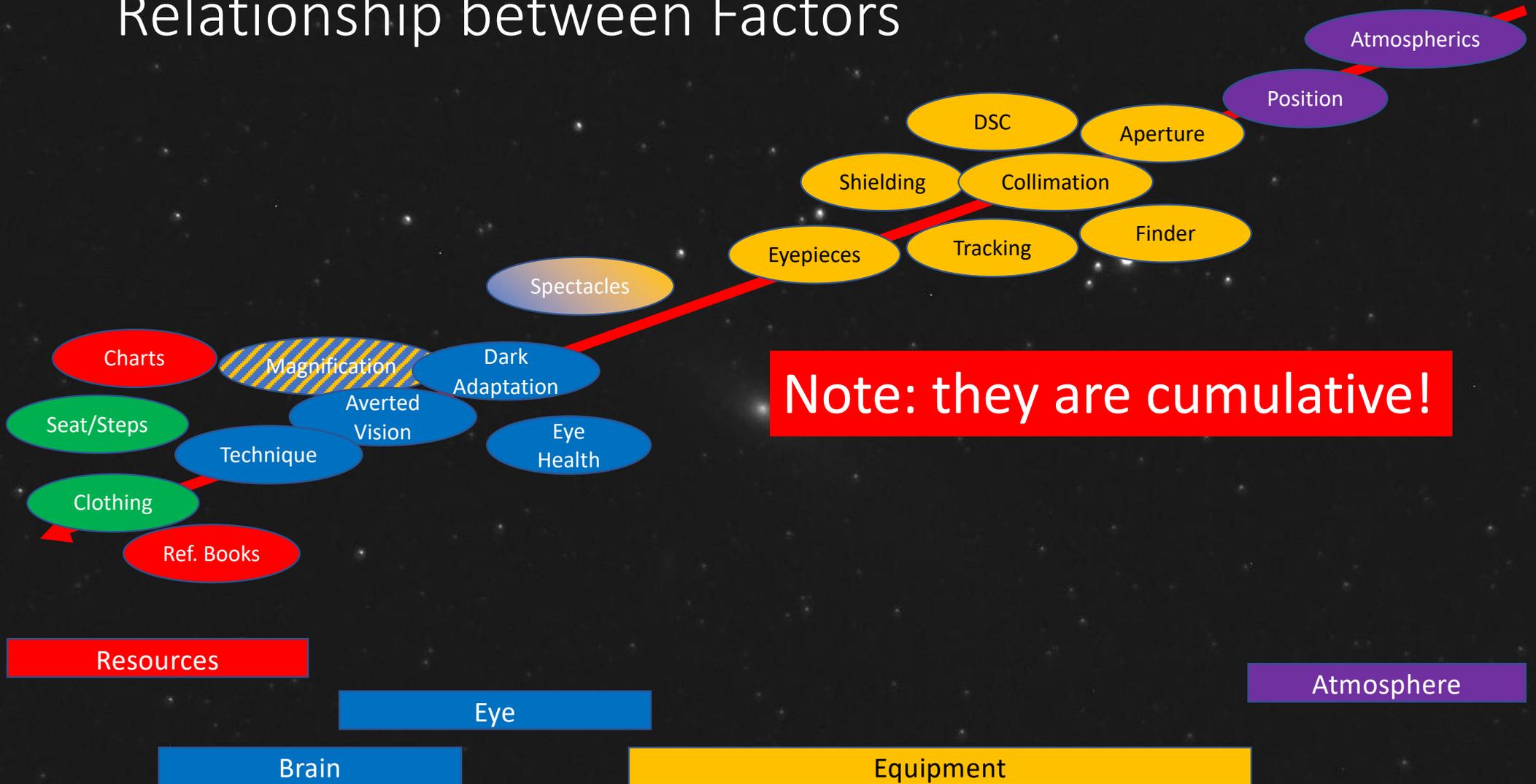
Less Difficult



More Difficult

- Equipment
- Ergonomics
- Vision
- The Sky
- Resources

# Relationship between Factors



# 5 top Factors in seeing difficult targets

Selection based on.....

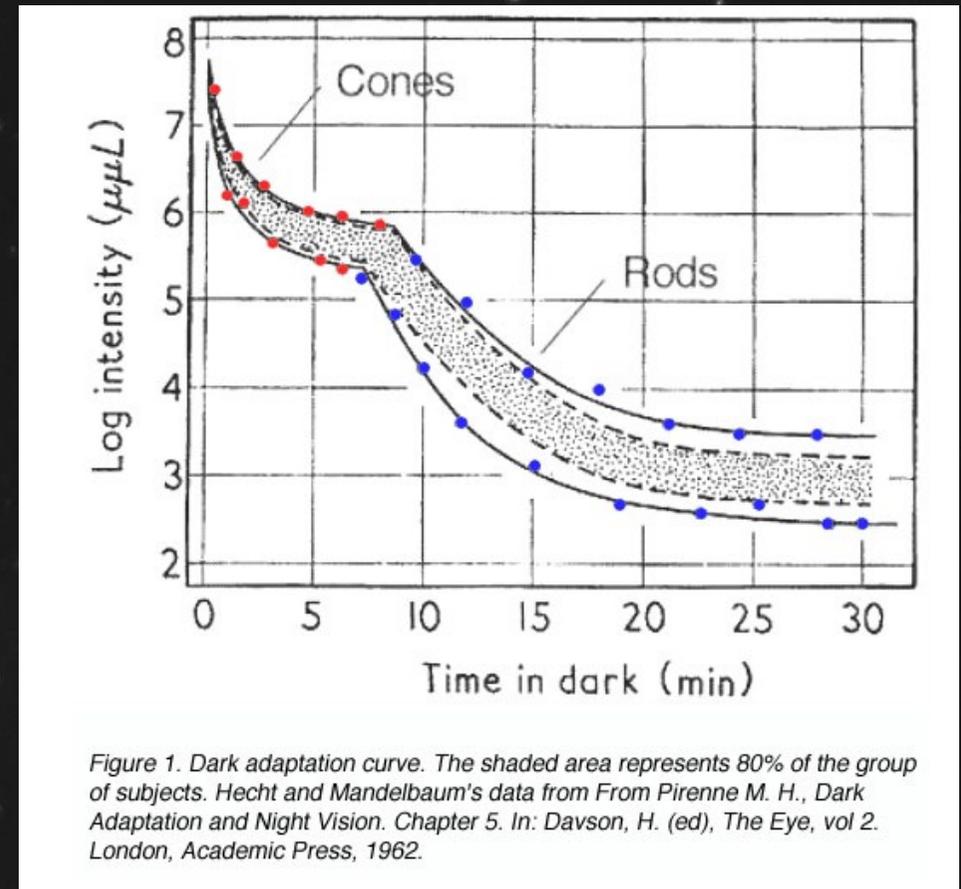
- 30 years experience
- Analysis of CN & SGL posts
- Survey
- 'Interviews'



- Achieving maximum Dark Adaptation
- Observing Technique
- Magnification & Exit Pupil
- Eyes
- Finders

# Dark Adaptation

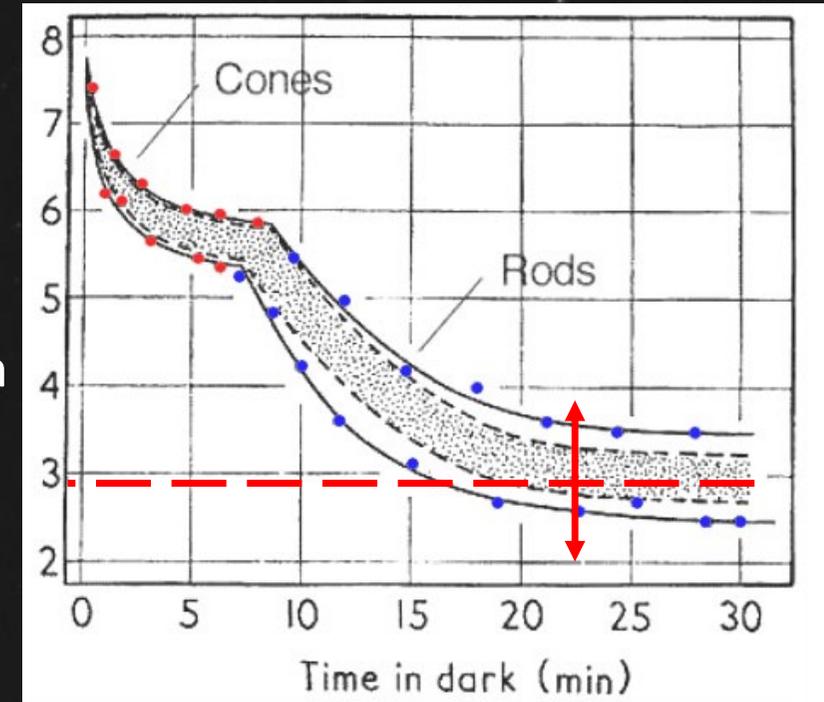
- The most important factor?
- Basics are well known
- Less well quantified
  - Impact of background levels
  - Actual impact of red light (dim and bright)
  - Recovery from various light exposures
- Issues
  - Wavelength of the red
  - Intensity of lights
  - Durations of adaptation & disruption
  - Distraction effects



Problem in 3-dimensional space: Colour - Intensity - Duration

## Dark Adaptation (cont.)

- Typical 'floor' of Dark Adaptation
  - ~ 2 mags on average higher than theoretical minimum
  - Anecdotal evidence of effect of different floors.
    - 6.5 is a 'typical' visual limiting magnitude
    - mag 8.5 if observing from a 'black box'
  - How to get from 6.5 to 8.5?
- Techniques to aid pre-adaptation
  - Use of red goggles, eye patches, etc (Sidgwick et al)
  - Eye Condition
- Experiments needed!

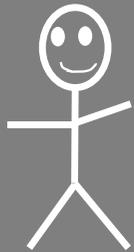


# Impact of ambient light level

Cepheus  
Altitude ~ 80deg



2 hours outside  
Sky was SQM21.5  
red torch  
Screens red and dim



18" f4.5  
x254

Limiting magnitude estimate  
15.0 (GAIA 2179697143977046784)

# Impact of ambient light level

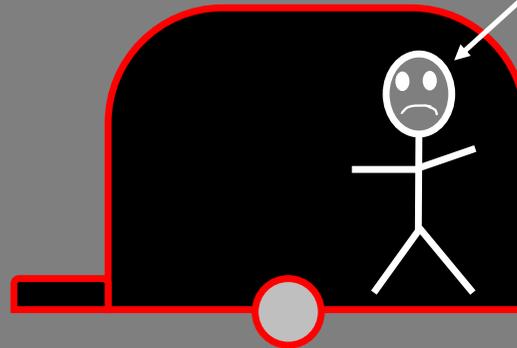
Cepheus  
Altitude ~ 80deg



18" f4.5  
x254

1 hour inside  
Total darkness  
No torches  
No screens etc

Me, bored!



# Impact of ambient light level

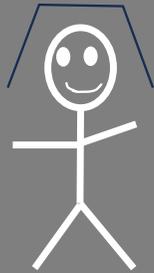
Cepheus  
Altitude  $\sim 80\text{deg}$



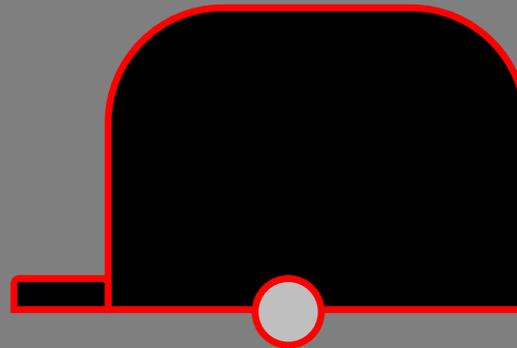
Immediately after

Limiting magnitude estimate  
16.8 (GAIA 2179696250623855360)

Hood



18" f4.5  
x254



1.8 mag improvement!

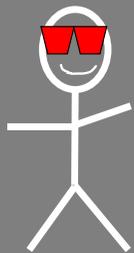
Close to the 'Ideal' limit  
of 17.0 (Clark)

# Impact of ambient light level

Cepheus  
Altitude  $\sim 80\text{deg}$



Continued outside  
Sky was SQM21.5  
Torches  
Screens



18" f4.5  
x254

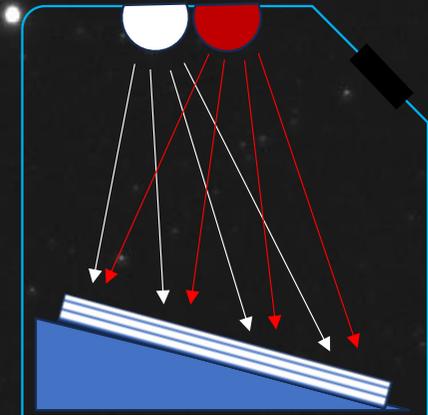
Limiting magnitude estimate  
 $\sim 16.1$

1 mag improvement!

close to the 'real' limit  
of 16 mag  
(Sinnott & Garstang)

# Another experiment

- Texas Star Party
  - Tested legibility of charts, using red & white light at various levels.
  - Measured impact on DA
- Dim white as good (better?) as bright red and much preferred for reading
- Recovery from dim white could be controlled.
  - 5-10 seconds of dim white reading, recovered full DA in ~30 seconds.
  - So, exposure to localised dim white light for short periods is OK (but problematic in a shared space)
- Recovery from red was faster
  - Dim red almost immediate, but charts couldn't be read satisfactorily
  - Brighter red took 15-20 seconds



# Red light

- Longer than 640nm avoids rod response. (Deep red)
  - Continuous exposure slows down Dark Adaptation a little
  - Fast recovery from strong exposure when already Dark Adapted.

= Deep red has almost no effect on DA

- But, difficult to read charts etc., unless rather bright

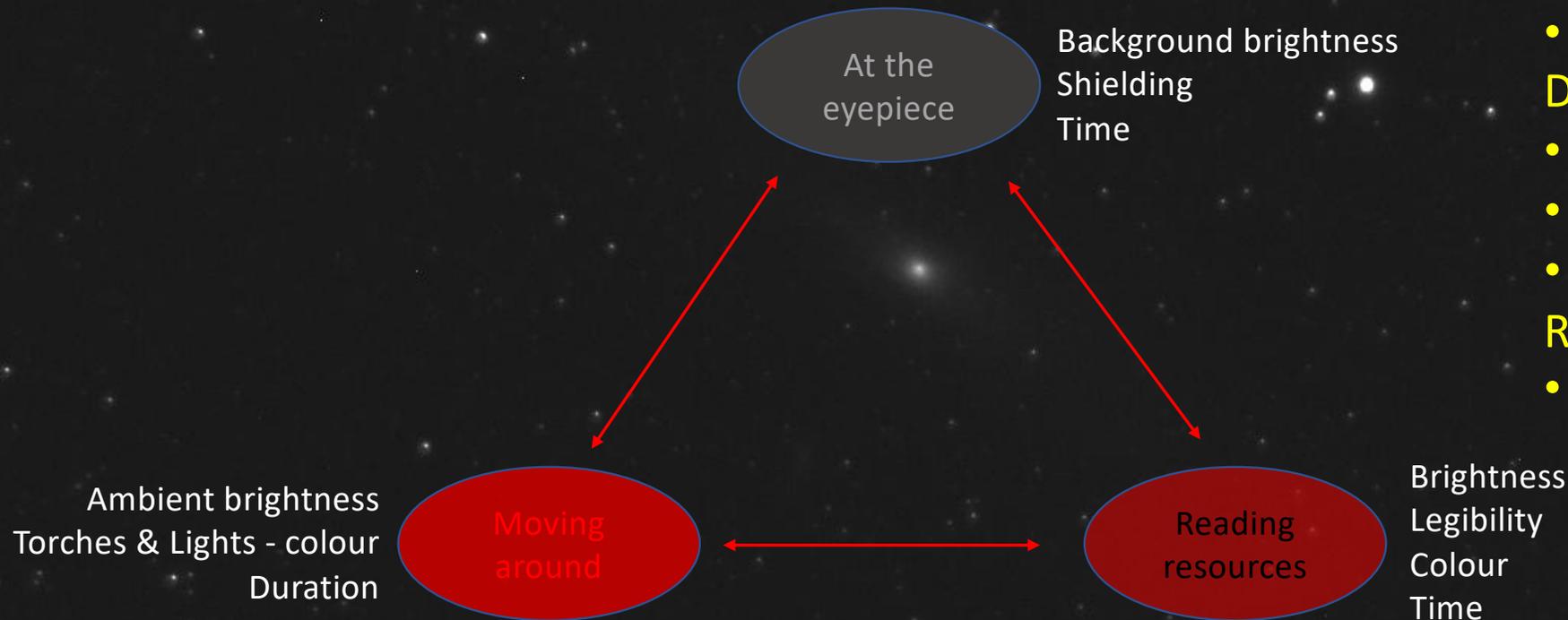
..... and can then be distracting

- Not a single astronomy red Led torch found > 630nm

# White light

- Continuous exposure limits degree of Dark Adaptation achievable
  - Sets a 'floor' to the Dark Adaptation curve.
  - Requires time at the eyepiece to lower the 'floor'. Can be at least 30 minutes.
  - Thus Red goggles can help achieve and maintain Dark Adaptation
- When Dark Adapted
  - Recovery to short exposure of dim white is quite rapid.
  - Strong exposure will bleach the retina according to the strength of the source and need some minutes to recover.
  - Small and intermittent dim white lights produce obvious distraction, but not loss of DA

# Managing your Dark Adaptation



## Prevent

- Dark site
- Rules
- Ruby Red light
- Shroud

## Deter

- Eye patches
- Goggles
- Hood

## Recover

- Time

# Averted Vision

- A technique to exploit Dark Adaptation
- Distribution of rods isn't uniform.
  - Sidgwick et al, upper - lower bias
  - Individual variability – requires personal exploration
- Dwell
  - Stare (6 second integration time)
  - Comfort (seat and clothing)
- Scope Tap
  - Supplements Averted Vision



Right-eye in object space

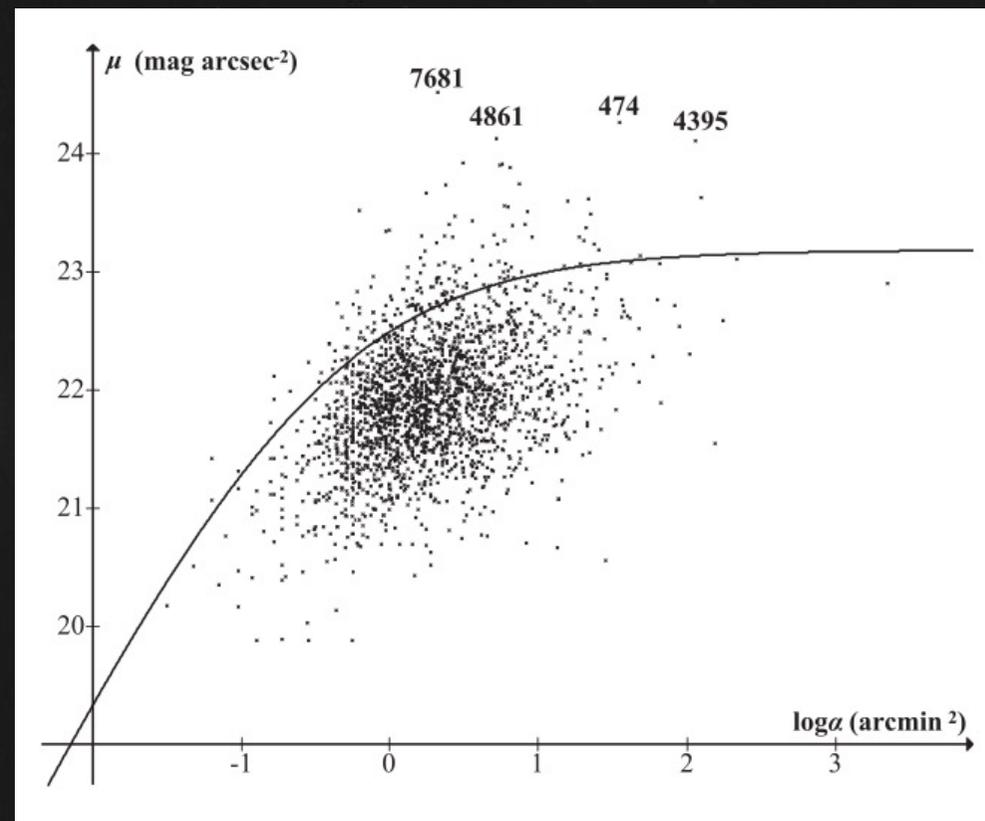
# Magnification & Exit Pupil

- Well known optimum magnification relationship to resolve fainter detail (ref, Roger N Clark et al)
  - Increasing magnification reduces background brightness – good for ‘point’ sources
    - 50% increase in magnification darkens background by about 1 mag.
  - Increasing magnification (larger apparent target) improves detection of contrast
    - But target is getting dimmer too.
- Nothing is straightforward!
  - The eye has a typical noise threshold of about 26 mpsas
  - WWII studies suggested large targets of 28-29 mpsas may be ‘detected’.
  - For dark skies and really low Surface Brightness targets, try reducing magnification.
- Anecdotal evidence that exit pupils <2mm can be counter-productive for some observers, and ~3mm may be an optimum.

# Human contrast threshold and astronomical visibility

Andrew Crumey, *MNRAS* 442, 2600–2619 (2014)

- Thorough re-examination of previous work (Blackwell, Curtiss, Clark, Schaefer, et al)
- Normalised previous disparate findings using a number of factors (Observer, equipment, light pollution, etc).
- Applied his new analysis to the established data. From Herschel to O'Meera
- Attributes Critical Visual Angle to the 'Ricco Area' effect (ref. Ricco's Law)
- Highlights 3 good sources that indicate a peak in visual acuity at 3mm exit pupil.



Herschel's 17.5'' ngc observations

# The Eyes & Eye Health

Eye Condition	Short term	Vitamin A good. Glucose & Oxygen for Rhodopsin regeneration.	<u>But.</u> quick changes in blood sugar can distort the eye.
	Long term	Diabetes impacts badly Cataracts Posterior Vitreous Detachment Macular Degeneration	L-lysine, Zinc, Vit C, Vit E, Lutein, Zeaxanthin
Spectacles		Eyepiece eye-relief issues Progressives Reduces eye-cup shielding benefits	Comfort Problematic
Alcohol & Smoking		Dims peripheral vision first Can cause fast change in blood sugar level – bad Decreases contrast sensitivity Slows pupil dilation	

# Atmospherics

- Air Mass effect
  - Not very significant
  - 0.1 mag at 45 deg
  - 1 mag at 10 deg
- Aerosols
  - Even on a good night
    - MODTRAN6 model, Mid-latitude Winter, 23km vis
    - Moisture 50% absorption at 45deg, 1 mag at 20deg
- Light Pollution
  - Sky brightening under LP is the biggest issue
  - Particles (dust etc) add very significantly to losses and SQM.

# 'Seeing' the Elusive – League table



**Remember – factors are additive**

Note: All the above are cumulative.

# Practical aspects of 'seeing' - 1

- 4 magnitude difference between rods and cones sensitivity
  - Must maintain and protect the fullest dark adaptation
    - More time at the eyepiece and much less moving around and at charts etc.
- Averted Vision
  - Most effective when the target position in the eyepiece field is known.
  - 2 magnitudes difference across the rods alone
    - Averted Vision is difficult enough to start with – searching the field is not effective
    - You can avert to your 'sweet spot', (and not your blind spot!)
    - You can employ the 6 sec dwell technique

# Practical aspects of 'seeing' - 2

- Let's assume a target dim enough to require steady averted vision

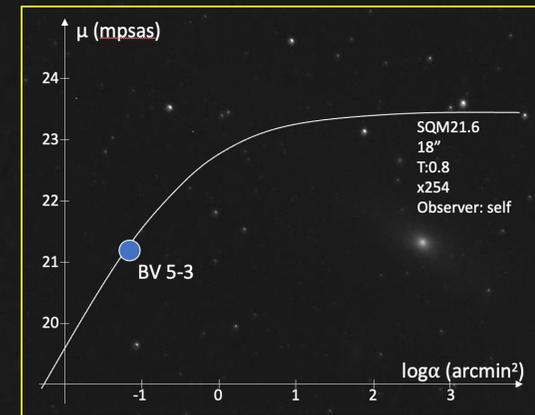
Do you,

- a. Get 'close enough' & visually search, or
- b. Spend more time locating and get spot on.

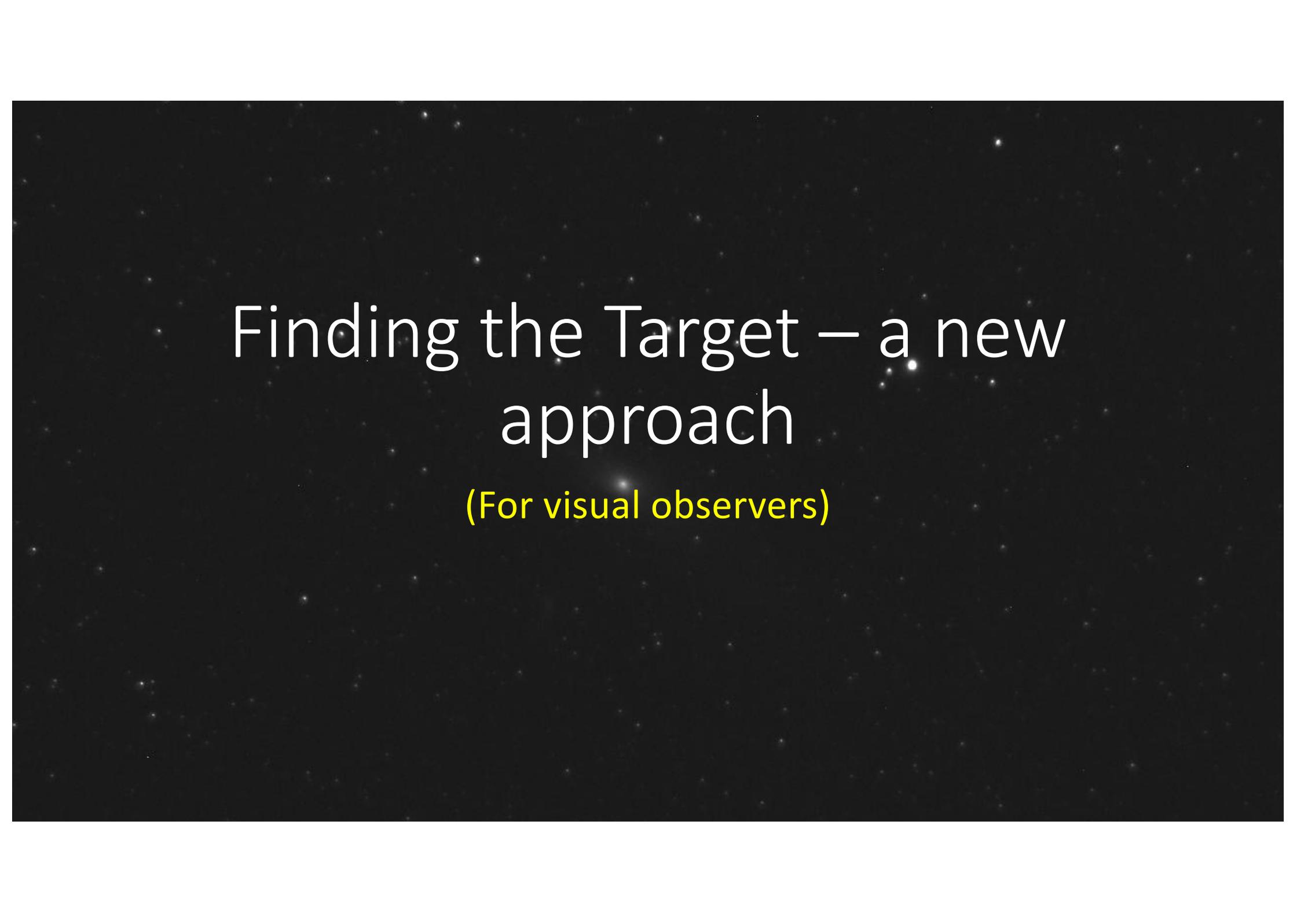
- Example: BV 5-3 - a challenging target (SB ~ 21.2mag)

- In the eyepiece field but 3 out of 4 observers couldn't see it.
- When told exactly where it was, 3 out of 4 could see it. (ie 2 more)

- Observing with a buddy often leads to seeing 'more'

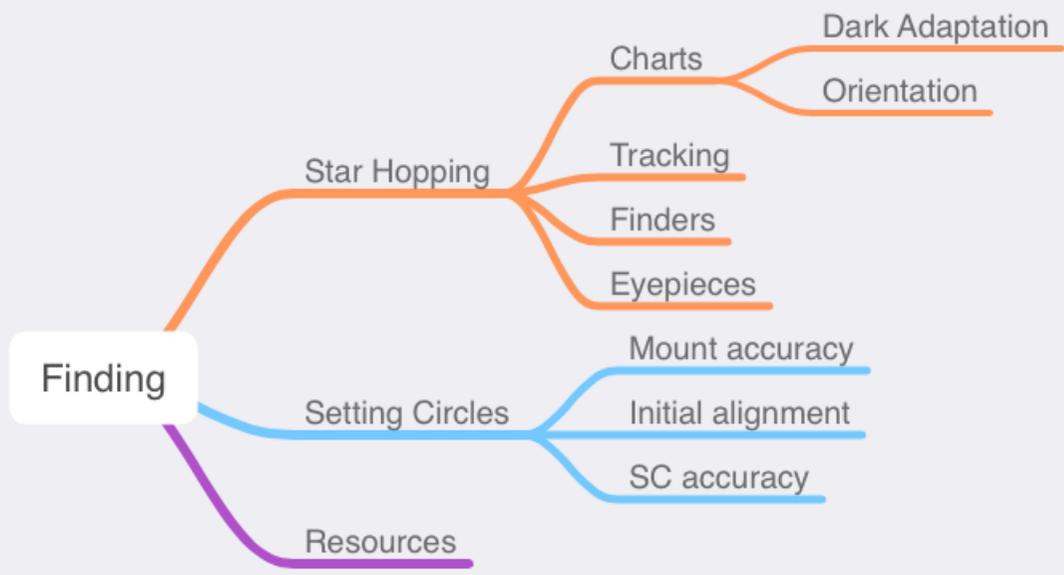


Knowing the position of the target in the field of view is critical



# Finding the Target – a new approach

(For visual observers)



# Finding the (elusive) Target

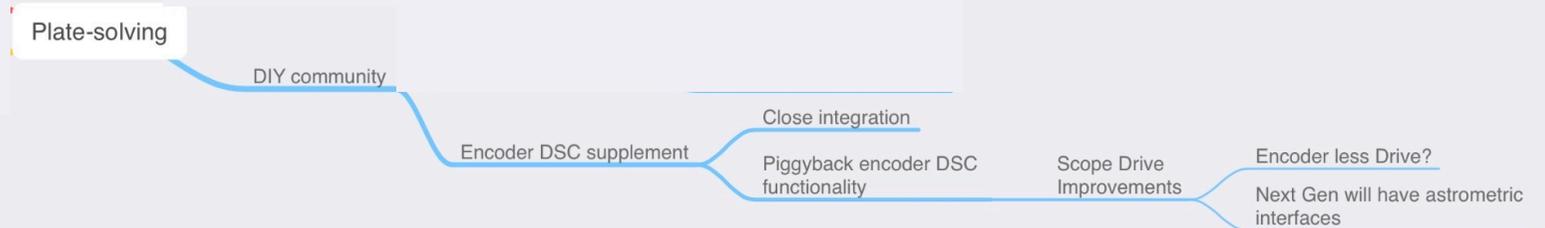
- Star-hopping
  - Low tech but requires practice and some skill
  - Maintaining Dark Adaptation while using charts needs care
  - Can take a time
  - Really benefits from a tracking mount
- Digital Setting Circles (DSC)
  - Relatively easy to fit
  - DSC in itself is accurate and powerful
  - Overall accuracy depends on the mount & initial alignment
    - Accuracy in practice is variable.
    - Usually still need some star-hopping (or 'local syncs') for difficult targets.
    - Many DSCs can build a model of a mount's systematic errors, but takes setup time
  - GoTo systems are dependent on DSC accuracy

# Past & Recent developments in Plate-Solving

- Astro-photographers have long moved to plate-solving
  - Small camera TFOV.
  - They want maximum time on target
  - They had cameras and computers already.
  - Plate-solving for most has completely removed the need for accurate mount alignment
- Many plate-solving applications run on small SBCs and even smart phones.
- Celestron StarSense is the first commercial use of a plate-solving finder for visual use.
  - Growing in popularity but not easily integrated with other scopes

# Future developments with Plate-Solving? (for Visual Observers)

Redo as commercial vs amateur diy  
Stand alone vs integrated



# 'eFinder' prototype



1 deg fov (custom f4 lens)  
Mechanical alignment needed  
LCD display issues  
Dedicated gps dongle needed  
Applies corrections to a DSC  
Excellent but slow solve performance

# 'eFinder'



6 deg fov (f1.8 cctv lens)  
Semi-automatic alignment  
Remote OLED display & control box  
Fully integrated with Nexus DSC  
Faster solving

# 'eFinder'

Alternative 'GUI' display  
on pc or tablet (via wifi)

Display can show eyepiece  
field of view

Can rotate to match the  
eyepiece view. Manual or  
automatic.

Any or all types of object  
can be automatically  
annotated.

Currently working on  
adding Comets

Writing an Android App to  
simplify its use.

Date 08 Sep 2022  
UTC 17:50:21  
LST 16h 58m 36s  
Loc: -0.71° 51.34°

Exposure    Gain  
● 0.1    ● 1  
● 0.5    ● 2  
● 1    ● 5  
● 5    ● 25  
● 10    ● 50

200mm finder  
 graticule  
 Polaris image  
 M13 Image

Offset: 0,0  
Measure  
Use New  
Save Offset  
Use Saved 12.6, 7.0  
Reset Offset

Align

	Nexus	delta x,y	Solution	delta x,y	Target
RA	16:56:32		16:42:41		
Dec	+35:46:27		+36:26:13		
Az	180:52:00	168.5	191:36:59		
Alt	+74:25:58	25.6	+74:51:34		

elapsed time 1.43 sec

Display  
 Scope centre  
 zoom x2  
 flip  
 mirror  
 auto-rotate  
 rotate angle 15

Annotate  
 Bright  
 Hip  
 H-D  
 ngc/ic  
 Abell  
 Tycho2

FOV indicator  
● 8mm Ethos  
● 10mm Ethos  
○ 13mm Ethos  
● 21mm Ethos

17:49:17 Nexus connected  
17:49:27 "/pixel: 3.75  
17:49:28 solved  
17:49:28 no GoTo target  
17:49:51 annotation successful

# Digital Finder features

- Display current telescope position, with respect to true RA & Dec
- Simplify initial 2 star alignment
  - Actual stars not needed!
- Perform 'Local Sync' anywhere, without an actual star alignment
- Refine a GoTo, (aka GoTo++)
- Simplify mount modelling (eg Mpoint)
- Derive motor drive rates without need for scope initial alignment - 'Point and track'



# Full integration is possible (and coming)

- Run the plate-solver code within the scope motor drive box.
- Plug a camera directly into the drive box
- Scope drive hand pad shows absolute scope eyepiece centre coordinates
- Accurate mount encoders are becoming less essential



# 'Seeking the (very) Elusive'

- Double check you are centered in the correct field
  - Finder Charts
  - OLED Tablet recommended
  - Consider a Digital Finder
- Maintain Dark Adaptation
  - Optimise your averted vision
- Be comfortable
  - clothing
  - seat
  - hydrated
  - fed
- Consider magnification & exit pupil
- Try nudging the scope
- Maintain good blood sugar levels – 'AstroBrownies'
- Reference images
  - eg, Night Sky Observers Guide
- Spend more time at the eyepiece!
- Ultra wide eyepieces may be a disadvantage.
- Experiment & practice

# In Conclusion.....

- With some effort, up to two magnitudes deeper is possible.
- No great expense, just method and diligence.
- A massive difference to the visual observing experience
- Digital Finders could simplify future scope drives and the user experience.
- They allow scopes to be pointed to sub-arc minute accuracy with respect to actual sky.
- Especially useful with compact and lighter mount designs
- For me, better DA & minimising the 'hunt' allowed much deeper observations

'Find'

'See'

'Look'

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